

## ON THE RIEMANNIAN PENROSE INEQUALITY WITH CHARGE AND THE COSMIC CENSORSHIP CONJECTURE

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ABSTRACT. We note an area-charge inequality originally due to Gibbons: if the outermost horizon  $S$  in an asymptotically flat electrovacuum initial data set is connected then  $|q| \leq r$ , where  $q$  is the total charge and  $r = \sqrt{A/4\pi}$  is the area radius of  $S$ . A consequence of this inequality, in conjunction with the positive mass theorem with charge, is that for connected black holes the following lower bound on the area holds:  $r \geq m - \sqrt{m^2 - q^2}$ . When combined with the upper bound  $r \leq m + \sqrt{m^2 - q^2}$  which is expected to hold always, this implies the natural generalization of the Riemannian Penrose inequality:  $m \geq 1/2(r + q^2/r)$ . We also establish the same lower bound without the assumption of time symmetry.

A natural generalization of the Riemannian Penrose inequality incorporating electric charge is:

$$(1) \quad m \geq \frac{1}{2} \left( r + \frac{q^2}{r} \right),$$

with equality if and only if the data is Reissner-Nordström. Here  $m$  is the ADM mass,  $r = \sqrt{A/16\pi}$  is the area radius of the outermost horizon, and  $q$  is the total charge. This inequality is known to hold when the outermost horizon is connected, but could be violated otherwise. The proof of the inequality for a connected horizon follows the inverse mean curvature flow argument of Huisken-Ilmanen [6], with a minor modification based on an earlier observation of Jang [8]. The rigidity statement was recently proved by Khuri and Disconzi [3]. The inequality (1) can fail when the outermost horizon is not connected, as shown in [10]. Indeed, the area of the cross-section of the necks in a Majumdar-Papapetrou solution with two bodies violates (1), and while this solution is not asymptotically flat, and does not contain a compact minimal surface, these deficiencies were corrected in [10] by gluing using a perturbation argument two identical copies along the necks. As observed already by Jang, inequality (1) is equivalent to the two inequalities:

$$(2) \quad m - \sqrt{m^2 - q^2} \leq r \leq m + \sqrt{m^2 - q^2}.$$

We note that  $|q| \leq m$  follows from the positive mass theorem with charge [5].

The upper bound on  $r$  in (2) is suggested by cosmic censorship, via a heuristic argument of Penrose. If the data violates this inequality, and if the evolution is smooth enough, then one expects the area of the horizon to be non-decreasing while the other parameters are constant. Hence the inequality will also be violated

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in the limit of late times, in contradiction to all the known stationary solutions without naked singularities. In a future paper, we prove this upper bound using a modification of the conformal flow used by H. Bray. An outline of the proof will appear in [7].

The counter-example constructed in [10] violates the lower bound in (2), and one might say that this is not so surprising since there is no physical motivation for this lower bound. On the other hand, as Robert Wald [9] pointed out, the fact that the lower bound is always satisfied when the outermost horizon is connected might be surprising, and indeed no physical motivation has so far been proposed for this inequality. In this short note, we show that in this case, the lower bound in fact follows from an inequality first proved by Gibbons [4] using the stability of the outermost horizon as an area minimizing surface. Thus, while Penrose's heuristic argument based on cosmic censorship provides a physical justification for the upper bound on  $r$ , our Corollary 1 below shows that the positive mass theorem with charge in conjunction with Gibbons' inequality provides a physical justification for the lower bound on  $r$  when the outermost horizon is connected.

We begin by very briefly introducing a few definitions. An initial data set  $(M, g, E)$  consists of a 3-manifold  $M$ , a Riemannian metric  $g$ , and a vector field  $E$ .<sup>†</sup> We assume that the data satisfies the Maxwell constraint  $\operatorname{div}_g E = 0$ , and the dominant energy condition  $R \geq 2|E|^2$ , where  $R$  is the scalar curvature of  $g$ . We assume that the data is strongly asymptotically flat meaning that there is a compact set  $K$  such that  $M \setminus K$  is the finite union of disjoint ends, and on each end the fields decay according to:

$$g - \delta = O(|x|^{-1}), \quad E = O(|x|^{-2}).$$

In addition we assume that  $R_g$  is integrable. This guarantees that the ADM mass and total charge

$$m = \frac{1}{16\pi} \int_{S_\infty} (g_{ij,j} - g_{jj,i}) \nu^i dA, \quad q = \frac{1}{4\pi} \int_{S_\infty} E_i \nu^i dA$$

are well defined. Here,  $\nu$  is the outer unit normal, and the limit is taken in a designated end. Conformally compactifying all but the designated end, we can now restrict our attention to surfaces which bound compact regions, and define  $S_2$  to enclose  $S_1$  to mean  $S_1 = \partial K_1$ ,  $S_2 = \partial K_2$  and  $K_1 \subset K_2$ . An *outermost horizon* is a compact minimal surface not enclosed in any other compact minimal surface.

A version of the following theorem was proved in [4, Section 6]. The main difference is that instead of stability we assume the outermost condition, which then in turn implies stability. The outermost condition is a natural condition appearing in statements of the Penrose inequality. We bring it here for completeness and because the proof is very simple.

**Theorem 1.** *Let  $(M, g, E)$  be strongly asymptotically flat, satisfying  $R \geq 2|E|^2$  and  $\operatorname{div}_g E = 0$ , and suppose the outermost horizon  $S$  is connected. Then*

$$|q| \leq r.$$

*Proof.* We begin by pointing out that  $S$  is in fact outer minimizing, meaning that it has area no greater than any other surface which encloses it, see for example [10,

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<sup>†</sup>For simplicity, we first assume that the magnetic field vanishes.

pg 706]. Thus,  $S$  is a stable minimal surface and from the second variation of area we get:

$$0 \leq \int_S -(|\chi|^2 + R_{\nu\nu}) dA = \int_S \left( \kappa - \frac{1}{2}|\chi|^2 - \frac{1}{2}R \right) dA,$$

where  $\chi$  is the second fundamental form of  $S$  in  $M$ ,  $\nu$  is the outward unit normal, and  $\kappa$  is the Gauss curvature of  $S$ . The second equality follows from the Gauss equation. Since  $S$  is connected, it follows from Gauss-Bonnet that  $\int_S \kappa dA = 4\pi$  hence

$$4\pi \geq \int_S \frac{1}{2}R dA \geq \int_S |E \cdot \nu|^2 dA \geq \frac{1}{4\pi r^2} \left( \int_S E \cdot \nu dA \right)^2 = \frac{(4\pi q)^2}{4\pi r^2}.$$

□

**Corollary 1.** *Under the same hypotheses as Theorem 1, we have:*

$$r \geq m - \sqrt{m^2 - q^2}.$$

*Proof.*

$$(3) \quad m = \sqrt{q^2 + m^2 - q^2} \leq |q| + \sqrt{m^2 - q^2} \leq r + \sqrt{m^2 - q^2}.$$

This proves the lower bound on  $r$ . □

Once the lower bound and the upper bound  $r \leq m + \sqrt{m^2 - q^2}$  both hold, inequality (1) follows. However, we point out that in fact we do not know of any independent proof of this upper bound.

We end by generalizing these observations in two ways, namely by including the magnetic field and removing the assumption of time symmetry. In this regard, consider the general initial data set  $(M, g, k, E, B)$ , where  $k$  is a symmetric 2-tensor representing the extrinsic curvature of  $M$  in spacetime, and  $B$  is a vector field representing the magnetic field. For strong asymptotic flatness, we require these quantities to satisfy the following fall-off conditions at spatial infinity

$$k = o(|x|^{-2}), \quad B = O(|x|^{-2}).$$

Similarly to the above, the total charges are now given by

$$q_e = \frac{1}{4\pi} \int_{S_\infty} E_i \nu^i dA, \quad q_b = \frac{1}{4\pi} \int_{S_\infty} B_i \nu^i dA,$$

and the matter and current densities for the non-electromagnetic matter fields are given by

$$\begin{aligned} 2\mu &= R + (\text{Tr } k)^2 - |k|^2 - 2(|E|^2 + |B|^2), \\ J &= \text{div}_g(k - (\text{Tr } k)g) + 2E \times B. \end{aligned}$$

The following is a Corollary of Theorem 2.1 in [2].

**Corollary 2.** *Let  $(M, g, k, E, B)$  be strongly asymptotically flat, satisfying  $\mu \geq |J|$  and  $\text{div}_g E = \text{div}_g B = 0$ , and suppose the outermost apparent horizon  $S$  is connected. Then*

$$(4) \quad r \geq m - \sqrt{m^2 - q_e^2 - q_b^2}.$$

*Proof.* As  $S$  is outermost, it is stable, and hence Theorem 2.1 in [2] implies the area-charge inequality

$$\sqrt{q_e^2 + q_b^2} \leq r.$$

Under the current assumptions the positive mass theorem with charge yields

$$m \geq \sqrt{q_e^2 + q_b^2}.$$

These two inequalities combine as in (3) to give

$$m \leq \sqrt{q_e^2 + q_b^2} + \sqrt{m - q_e^2 - q_b^2} \leq r + \sqrt{m - q_e^2 - q_b^2}$$

and (4) follows.  $\square$

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#### REFERENCES

1. H. BRAY, *Proof of the Riemannian Penrose conjecture using the positive mass theorem*, J. Differential Geom. **59** (2001), 177–267.
2. S. DAIN, J. JARAMILLO, AND M. REIRIS, *Area-charge inequality for black holes*, Class. Quantum Grav. **29** (2012), p. 035013
3. M. M. DISCONZI AND M. A. KHURI, *On the Penrose inequality for charged black holes*, Class. Quantum Grav. **29** (2012) arxiv.org/abs/1207.5484, p. 245019.
4. G. W. GIBBONS, *Some comments on gravitational entropy and the inverse mean curvature flow*, Classical Quantum Gravity **16** (1999), no. 6, 1677–1687.
5. G. W. GIBBONS, S. HAWKING, G. HOROWITZ AND M. PERRY, *Positive mass theorem for black holes*, Commun. Math. Phys. **88** (1983), 295–308.
6. G. HUISKEN AND T. ILMANEN, *The Inverse Mean Curvature Flow and the Riemannian Penrose Inequality*, J. Differential Geom. **59** (2001), 353–437.
7. M. KHURI, G. WEINSTEIN AND S. YAMADA, *The Riemannian Penrose inequality with charge for multiple black holes*, in preparation.
8. P. S. JANG, *Note on Cosmic Censorship*, Phys. Rev. D. **20** (1979) No. 4, 834–838.
9. R. M. WALD, Private communication.
10. G. WEINSTEIN AND S. YAMADA, *On a Penrose Inequality with Charge*, Comm. Math. Phys. **257** (2005) No. 3, 703–723.

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