

## Remarks on Scattering on Scattering Manifolds

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### Abstract

In this talk, we discuss scattering theory for a class of manifolds. We consider the asymptotic completeness and the microlocal properties of the scattering matrix. The space we consider is called *scattering manifolds* following R. Melrose, and we construct a time-dependent scattering theory for Schrödinger operators on such manifolds. In particular, we discuss an alternative approach to a theorem by R. Melrose and M. Zworski on the microlocal properties of the *absolute scattering matrix*. This work is partly in progress, and several theorems are preliminary.

**Model:** We consider an  $n$ -dimensional noncompact manifold (without boundary):

$$M = M_0 \cup M_\infty$$

where  $M_0$  is relative compact, and  $M_\infty$  is diffeomorphic to  $(1, \infty) \times \partial M$ , where  $\partial M$  is a closed manifold without boundary. We consider  $\partial M$  as a boundary of  $M$  at infinity. We fix an identification map:

$$I : M_\infty \cong (1, \infty) \times \partial M \ni (r, \theta), \quad r \in (1, \infty), \theta \in \partial M.$$

Let  $g^\partial$  be a Riemannian metric on  $\partial M$ , and we denote

$$g^\partial = \sum_{i,j} g_{ij}^\partial(\theta) d\theta^i d\theta^j, \quad \theta \in \partial M.$$

**Definition:** A Riemannian metric  $g^{cn}$  on  $M$  is called *conic* if it has the following form:

$$g^{cn} = dr^2 + r^2 g^\partial \quad \text{on } M_\infty,$$

where we identify  $M_\infty$  with  $(1, \infty) \times \partial M$  as above.

**Example 0:** (Euclidean space)  $M = \mathbb{R}^n$ ,  $\partial M = S^{n-1}$ ,  $g^\partial = d\theta^2$  is the surface metric on  $S^{n-1}$ . Then  $g^{cn} = dr^2 + r^2 d\theta^2$  is the standard flat metric on  $\mathbb{R}^n$  in the polar coordinate on  $M_\infty = \{x \mid |x| > 1\}$ . The identification map is

$$I : r\theta \in M_\infty \mapsto (r, \theta) \in (1, \infty) \times S^{n-1}.$$

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This is a typical example and we should keep this in mind in the following argument.

**Example 1:** (Conic metric on  $\mathbb{R}^n$ ) Let  $M$  and  $\partial M$  as in Example 0, but we introduce a different metric on  $S^{n-1}$ . Then we have a different conic metric structure on  $\mathbb{R}^n$ . For example, we can set  $g^\partial = \alpha d\theta^2$  with  $\alpha > 0$ , and we have a different geometric structure.

**Definition:** A Riemannian metric  $g$  on  $M$  is called *scattering metric* if

$$g = g^{cn} + m,$$

where  $g^{cn}$  is the conic metric, and  $m$  is a symmetric 2-form such that

$$m = m^0(r, \theta) dr^2 + r \sum_{j=1}^{n-1} m_j^1(r, \theta) (dr d\theta^j + d\theta^j dr) + r^2 \sum_{i,j=1}^{n-1} m_{ij}^2(r, \theta) d\theta^i d\theta^j$$

on  $M_\infty$ , and the coefficients satisfy

$$|\partial_r^k \partial_\theta^\alpha m_*^\ell(r, \theta)| \leq C_{k\alpha} r^{-\mu_\ell - k}, \quad (r, \theta) \in M_\infty$$

for any  $k, \alpha, \ell = 0, 1, 2$  with  $\mu_\ell > 0$ .

Scattering metric was defined originally by R. Melrose [2], but here we use an equivalent, but different definition. (This formulation was introduced in [1]). We will assume the metric perturbation  $m$  is short-range type in the following sense:

**Definition:** A metric  $g$  on  $M$  is called *short-range type* if

$$\mu_0 > 1, \quad \mu_1 > 1/2, \quad \mu_2 > 0.$$

Let  $\Delta_g$  be the Laplace-Beltrami operator on  $M$  corresponding to the Riemannian metric, i.e.,

$$\Delta_g = \frac{1}{\sqrt{G(x)}} \sum_{j,k=1}^n \partial_{x_j} g^{jk}(x) \sqrt{G(x)} \partial_{x_k}$$

where  $G(x) = \det(g_{jk}(x))$  and  $(g^{jk}) = (g_{jk})^{-1}$ .

**Definition:** A potential function  $V \in C^\infty(M; \mathbb{R})$  is called *short-range type* if there is  $\mu_3 > 1$  such that for any  $\alpha$  and  $k$ ,

$$|\partial_r^k \partial_\theta^\alpha V(r, \theta)| \leq C_{k\alpha} r^{-\mu_3 - k}, \quad (r, \theta) \in M_\infty.$$

In the following, we assume  $g$  and  $V$  are short-range type. We set

$$H = -\Delta_g + V(x) \quad \text{on } \mathcal{H} = L^2(M, \sqrt{G} dx).$$

**Proposition 1.**  $H$  is essentially self-adjoint on  $C_0^\infty(M)$ . Moreover,  $\sigma_{ess}(H) = [0, \infty)$ ;  $\sigma_p(H)$  is discrete with possible accumulation points only at 0;  $\sigma_{sc}(H) = \emptyset$ ; and  $\sigma_{ac}(H) = [0, \infty)$ .

*Idea of Proof.* Let  $j(r) \in C^\infty(\mathbb{R})$  be a smooth cut-off function such that

$$j(r) = \begin{cases} 1 & (r \geq 3/2), \\ 0 & (r \leq 1). \end{cases}$$

We use the Mourre theory with the conjugate operator:

$$A = \frac{1}{2i} \left( j(r)r \frac{\partial}{\partial r} + \frac{\partial}{\partial r} j(r)r + \frac{1}{2} j(r)r \partial_r(\log G(x)) \right)$$

on  $M^\infty$ . Then the rest of the argument is similar to the Euclidean case.  $\square$

**Scattering Theory:** We first construct a *free system*. We might use  $-\Delta_{g^{cn}}$  as the free system, but this operator itself is not very easy to handle. So, instead, we set

$$H_{fr} = -\frac{\partial^2}{\partial r^2} \quad \text{on} \quad M_{fr} = \mathbb{R} \times \partial M,$$

$$\mathcal{H}_{fr} = L^2(M_{fr}, dr \cdot \sqrt{g^\partial} d\theta)$$

$$J : \mathcal{H}_{fr} \rightarrow \mathcal{H}, \quad \text{where} \quad J\varphi(r, \theta) = \begin{cases} 0 & \text{on } M_\infty^c \\ j(r) (\det g^\partial(\theta)/G(r, \theta))^{1/4} \varphi(r, \theta) & \text{on } M_\infty. \end{cases}$$

$J$  is defined so that  $J$  is isometry on  $L^2([3/2, \infty) \times \partial M)$ . Note, asymptotically,  $J\varphi \sim r^{-(n-1)/2} \varphi$  as  $r \rightarrow \infty$ . In fact, for Examples 0 and 1, we have

$$J\varphi(r, \theta) = j(r)r^{-(n-1)/2} \varphi(r, \theta) \quad \text{for } \varphi \in L^2(\mathbb{R} \times S^{n-1}).$$

In this case, if we set  $\varphi = e^{ikr}$ , a generalized eigenfunction of  $H_{fr}$ , then

$$J\varphi = j(r)r^{-(n-1)/2} e^{ikr},$$

which is a spherical wave (generalized eigenfunction of  $\Delta$  for large  $r$ ).

We then set the wave operators:

$$W_\pm := \text{s-lim}_{t \rightarrow \pm\infty} e^{itH} J e^{-itH_{fr}} : \mathcal{H}_{fr} \rightarrow \mathcal{H}.$$

The existence of  $W_\pm$  is easy to show by the standard Cook-Kuroda method. We note  $g^{jk}$  has the form:

$$(g^{ij}) = (g_{ij})^{-1} = \begin{pmatrix} 1 + a_0 & r^{-1} a_1^t \\ r^{-1} a_1 & r^{-2} g_\partial + r^{-2} a_2 \end{pmatrix}$$

in the  $(r, \theta)$  coordinate, where  $\partial_r^k \partial_\theta^\alpha a_0 = O(r^{-1-\mu-k})$  and  $\partial_r^k \partial_\theta^\alpha a_j = O(r^{-\mu-k})$  for  $j = 1, 2$  with some  $\mu > 0$ . Here we denote  $g_\partial = (g^\partial)^{-1}$ .

We then set

$$\mathcal{H}_{fr, \pm} = \{ \varphi \in \mathcal{H}_{fr} \mid \text{supp } \hat{\varphi} \subset \mathbb{R}_\pm \times \partial M \},$$

where  $\hat{\varphi}$  is the Fourier transform of  $\varphi$  in  $r$ , i.e.,

$$\hat{\varphi}(\rho, \theta) = (\mathcal{F}\varphi)(\rho, \theta) = \int_{-\infty}^{\infty} e^{-i\rho r} \varphi(r, \theta) dr.$$

Then it is not difficult to see by the stationary phase method that

$$W_{\pm}(\mathcal{H}_{f_{r,\mp}}) = 0,$$

and hence it is natural to consider  $W_{\pm} : \mathcal{H}_{f_{r,\pm}} \rightarrow \mathcal{H}$ .

**Theorem 2.**  $W_{\pm}$  are isometry from  $\mathcal{H}_{f_{r,\pm}}$  to  $\mathcal{H}$ , and they are complete, i.e.,  $\text{Ran } W_{\pm} = \mathcal{H}_c(H)$ . Hence, in particular, the scattering operator defined by

$$S = W_{+}^{*}W_{-} : \mathcal{H}_{f_{r,-}} \rightarrow \mathcal{H}_{f_{r,+}}$$

is unitary.

*Idea of Proof.* Let

$$\begin{aligned} H_{\partial} &= -\frac{1}{\sqrt{G(x)}} \sum_{j,k=1}^{n-1} \partial_{\theta_j} j(r) g_{\partial}^{jk}(\theta) \sqrt{G(x)} \partial_{\theta_k} \\ &= -\sum_{j,k=1}^{n-1} \partial_{\theta_j} j(r) g_{\partial}^{jk}(\theta) \partial_{\theta_k} + (\text{lower order terms}). \end{aligned}$$

This operator is, roughly speaking, the pull-back of the Laplace-Beltrami operator on  $\partial M$  to  $M$ . By the Mourre theory, we can show

$$\langle j(r)r \rangle^{-\alpha} (H - \lambda \pm 0)^{-1} \langle j(r)r \rangle^{-\alpha} \in B(\mathcal{H}), \quad \lambda \in \mathbb{R}_{+} \setminus \sigma_p(H), \alpha > 1/2,$$

but these are not sufficient to show the completeness, since perturbation terms:  $r^{-1}a_1\partial_r\partial_{\theta}$ ,  $r^{-2}a_2\partial_{\theta}\partial_{\theta}$  are only of  $O(r^{-\mu})$ ,  $\mu > 0$ , with respect to  $H$ . Instead, we show

$$\langle j(r)r \rangle^{-\alpha} (H_{\partial} + 1)(H - \lambda \pm i0)^{-1} (H_{\partial} + 1)^{-1} \langle j(r)r \rangle^{-\alpha} \in B(\mathcal{H}), \quad \lambda \in \mathbb{R}_{+} \setminus \sigma_p(H).$$

These estimates are proved by resolvent equations and commutator computations. These imply that

$$(H - \lambda \pm i0)^{-1} : (H_{\partial} + 1)^{-1} \langle j(r)r \rangle^{-\alpha} \mathcal{H} \mapsto (H_{\partial} + 1)^{-1} \langle j(r)r \rangle^{\alpha} \mathcal{H}$$

is bounded, and this is sufficient to show the completeness by using the abstract stationary scattering theory.  $\square$

**Scattering Matrix:** By the intertwining property, we have

$$H_{f_r} S = S H_{f_r} : \mathcal{H}_{f_{r,-}} \rightarrow \mathcal{H}_{f_{r,+}},$$

and hence

$$\rho^2(\mathcal{F}S\mathcal{F}^{-1}) = (\mathcal{F}S\mathcal{F}^{-1})\rho^2 : \hat{\mathcal{H}}_{fr,-} \rightarrow \hat{\mathcal{H}}_{fr,+},$$

where  $\hat{\mathcal{H}}_{fr,\pm} = L^2(\mathbb{R}_\pm \times \partial M)$ . Thus,  $(\mathcal{F}S\mathcal{F}^{-1})$  commutes with multiplication by functions of  $\rho$ , and then we learn that  $(\mathcal{F}S\mathcal{F}^{-1})$  is decomposed as

$$(\mathcal{F}S\mathcal{F}^{-1})\varphi(\rho, \theta) = (S(\rho)\varphi(-\rho))(\theta), \quad \rho > 0, \varphi \in \hat{\mathcal{H}}_{fr,+}$$

with  $S(\rho) : L^2(\partial M) \rightarrow L^2(\partial M)$ , unitary.  $S(\rho)$  is called the *scattering matrix*.

**Melrose-Zworski Theorem:** Let

$$h(\theta, \omega) = \sum_{j,k} g_\partial^{jk}(\theta) \omega_j \omega_k \quad \text{for } (\theta, \omega) \in T^*\partial M$$

be the classical Hamiltonian on  $\partial M$ , and let  $\exp tH_{\sqrt{h}}$  be the Hamilton flow generated by  $\sqrt{h}$ , which is in fact the geodesic flow. Then we can show

**Theorem 3.**  $S(\rho)$  is an FIO corresponding to the canonical transform  $\exp \pi H_{\sqrt{h}}$ . In particular,

$$WF(S(\rho)\varphi) = \exp \pi H_{\sqrt{h}}(WF(\varphi)), \quad \varphi \in L^2(\partial M).$$

This result is a generalization of a result by R. Melrose and M. Zworski [3], though they used different definition of the scattering matrix, which is called the *absolute scattering matrix*. The absolute scattering matrix is defined as follows: Let  $\psi$  be a generalized eigenfunction of  $H$ :  $H\psi = \rho^2\psi$ . Then  $\psi$  has an asymptotic form:

$$\psi(r, \theta) \sim r^{-(n-1)/2} (e^{ir\rho} \varphi_+(\theta) + e^{-ir\rho} \varphi_-(\theta)) \quad \text{as } r \rightarrow \infty$$

with some  $\varphi_\pm \in L^2(\partial M)$ . The map:

$$\tilde{S}(\rho) : \varphi_+ \mapsto \varphi_-$$

is well-defined and  $\tilde{S}(\rho)$  is called the absolute scattering matrix since it is defined without using the time-dependent scattering theory. However, we can show

$$\tilde{S}(\rho) = -S(\rho)^{-1}$$

in our notation. As well as the formulation, the proof of Theorem 3 is considerably different from the one by Melrose and Zworski.

**Example 0:** (revisited) For the Euclidean case,  $\exp \pi H_{\sqrt{h}}(\theta, \omega) = (-\theta, -\omega)$ . Hence, the singularity of  $\varphi$  is mapped by the scattering matrix to the anti-podal points, which is well-known.

**Example 1:** (revisited) Let  $n = 2$  and we set  $g_\partial = \alpha g_0$  with  $\alpha > 0$  and  $g_0 = d\theta^2$ , the standard length on  $S^1$ . Then  $S(\rho)$  has a different microlocal propagation properties. Namely,

$$WF(S(\rho)\varphi) \subset \{\theta \pm \alpha\pi \mid \theta \in WF(\varphi)\}.$$

**Classical Scattering Theory for Conic Metric:** In order to understand the meaning of the Melrose-Zworski theorem, let us consider the classical scattering for the conic metric. Let  $p(r, \theta, \rho, \omega)$  be the classical Hamiltonian for the conic metric:

$$p(r, \theta, \rho, \omega) = \rho^2 + \frac{1}{\rho^2} \sum_{j,k} g_{\partial}^{jk}(\theta) \omega_j \omega_k, \quad r > 0, \rho \in \mathbb{R}, (\theta, \omega) \in T^* \partial M.$$

Let  $(r(t), \theta(t), \rho(t), \omega(t))$  be the solution to the Hamiltonian equation:

$$\dot{r} = \frac{\partial p}{\partial \rho}, \quad \dot{\theta} = \frac{\partial p}{\partial \omega}, \quad \dot{\rho} = -\frac{\partial p}{\partial r}, \quad \dot{\omega} = -\frac{\partial p}{\partial \theta},$$

with  $r(0) = r_0$ ,  $\theta(0) = \theta_0$ , etc. It is easy to see  $h(\theta(t), \omega(t))$  is invariant, i.e.,  $h(\theta(t), \omega(t)) = h(\theta_0, \omega_0) = h_0$ . Then we can solve equation for  $(r, \rho)$  easily to obtain

$$r(t) = \sqrt{4E_0 t^2 + 4r_0 \rho_0 t + r_0^2}, \quad E_0 = p(r_0, \theta_0, \rho_0, \omega_0),$$

and  $(\theta, \omega)$  satisfies the equation:

$$\dot{\theta} = \frac{1}{r^2} \frac{\partial h}{\partial \omega}, \quad \dot{\omega} = -\frac{1}{r^2} \frac{\partial h}{\partial \theta}.$$

So, by changing the time variable  $t \mapsto \tau(t) = \int_0^t ds / r(s)^2$ , we have

$$(\theta(t), \omega(t)) = \exp(\tau(t) H_h)(\theta_0, \omega_0),$$

where  $\exp(t H_h)$  is the Hamilton flow generated by  $h$  on  $T^* \partial M$ . As  $t \rightarrow \pm\infty$ ,  $\tau(t)$  converges to finite values:

$$\lim_{t \rightarrow \pm\infty} \tau(t) = \tau_{\pm} = \frac{1}{2\sqrt{h_0}} \left( \pm \frac{\pi}{2} - \tan^{-1} \frac{r_0 \rho_0}{\sqrt{h_0}} \right).$$

Hence we have

$$\lim_{t \rightarrow \pm\infty} (\theta(t), \omega(t)) = (\theta_{\pm}, \omega_{\pm}) = \exp(\tau_{\pm} H_h)(\theta_0, \omega_0).$$

Similarly, we can show by straightforward computations,

$$\lim_{t \rightarrow \pm\infty} \rho(t) = \rho_{\pm} = \pm \sqrt{E_0}, \quad \lim_{t \rightarrow \pm\infty} (r(t) - 2t\rho(t)) = r_{\pm} = \pm \frac{r_0 \rho_0}{\sqrt{E_0}}$$

This gives us the explicit formula for the (inverse) classical scattering operator:

$$(W_{\pm}^{cl})^{-1} : (r_0, \rho_0, \theta_0, \omega_0) \mapsto (t_{\pm}, \rho_{\pm}, \theta_{\pm}, \omega_{\pm}) = \lim_{t \rightarrow \pm\infty} (r(t) - 2t\rho(t), \rho(t), \theta(t), \omega(t)).$$

We note that the corresponding free Hamiltonian is simply given by  $\rho^2$ , which generates the free motion :  $(r, \rho, \theta, \omega) \mapsto (r + 2\rho t, \rho, \theta, \omega)$ . By the formula, it is easy to show

$$(W_{\pm}^{cl})^{-1} : (\mathbb{R}_+ \times \mathbb{R}) \times (T^* \partial M) \rightarrow (\mathbb{R} \times \mathbb{R}_{\pm}) \times (T^* \partial M)$$

is diffeomorphic, and hence

$$S^{cl} = (W_+^{cl})^{-1} \circ W_-^{cl} : (\mathbb{R}_- \times \mathbb{R}) \times (T^*\partial M) \rightarrow (\mathbb{R}_+ \times \mathbb{R}) \times (T^*\partial M)$$

is also diffeomorphic. In fact, we can easily show

$$S^{cl} : (r, \rho, \theta, \omega) \mapsto (-r, -\rho, \exp((\tau_+ - \tau_-)H_h)(\theta, \omega)),$$

and  $\tau_+ - \tau_- = \pi/(2\sqrt{\hbar_0})$ . In general, we have  $\exp(tH_q) = \exp((2t\sqrt{q})H_{\sqrt{q}})$  for  $q \geq 0$ , and hence we learn

$$\exp((\tau_+ - \tau_-)H_h)(\theta, \omega) = \exp(\pi H_{\sqrt{\hbar}}).$$

Thus we have

$$S^{cl} = (-I) \otimes \exp(\pi H_{\sqrt{\hbar}}),$$

and we realize that the Melrose-Zworski theorem is a quantization of this observation.

**Scattering Calculus:** In the proof of Theorem 3, we use the *scattering calculus* following Melrose [2], but again in a quite different formulation. For  $a \in C_0^\infty(T^*(\mathbb{R}_+ \times \partial M))$  (or  $\in C_0^\infty(T^*(\mathbb{R} \times \partial M))$ ), we denote the *scattering quantization* by

$$A = a(\hbar r, \theta, D_r, \hbar D_\theta), \quad \hbar > 0.$$

Note the difference of the location of the semiclassical parameter  $\hbar > 0$  from the usual semiclassical quantization  $a(r, \theta, \hbar D_r, \hbar D_\theta)$ . We identify  $\mathbb{R}_+ \times \partial M$  with  $M_\infty$ , and we consider  $A$  as an operator on  $L^2(M, \sqrt{G} dx)$ . For such an operator  $A$ , we consider

$$A(t) = e^{itH_{fr}} J^* e^{-itH} A e^{itH} J e^{-itH_{fr}}, \quad t \in \mathbb{R}.$$

$A(t)$  satisfies the Heisenberg equation:

$$\frac{d}{dt} A(t) = i[T(t), A(t)] + (\text{lower order error terms})$$

where

$$T(t) = e^{itH_{fr}} (HJ - JH_{fr}) e^{-itH_{fr}} \sim j(r - 2tD_r) \cdot \frac{h(\theta, D_\theta)}{(r - 2tD_r)^2}$$

as  $r \rightarrow \infty$ . We can construct the asymptotic solution to the Heisenberg equation:

$$A(t) = b_\hbar^t(\hbar r, \theta, D_r, \hbar D_\theta) \quad \text{where} \quad b_\hbar^t \sim b_0(\hbar^{-1}t, r, \theta, \rho, \omega) + O(\hbar),$$

and  $b_0$  can be computed explicitly using the classical flow. We let  $t \rightarrow \pm\infty$  and we learn

$$\lim_{t \rightarrow \pm\infty} A(t) = W_\pm^* A W_\pm \sim b_\hbar^\pm(\hbar r, \theta, D_r, \hbar D_\theta),$$

where  $b_\hbar^\pm \sim (a \circ W_\pm^{cl})(r, \theta, \rho, \omega) + O(\hbar)$ . Using this procedure again, we learn

$$SAS^{-1} = c_\hbar(\hbar r, \theta, D_r, \hbar D_\theta), \quad \text{where} \quad c_\hbar \sim a \circ (S^{cl})^{-1} + O(\hbar).$$

If  $A = a_1(D_r) a_2(\theta, \hbar D_t \hbar)$ , then we learn

$$SAS^{-1} \sim a_1(-D_r)(a_2 \circ \exp(\pi H_{\sqrt{\hbar}}))(\theta, D_\theta),$$

and hence

$$S(\rho)a_x(\theta, \hbar D_\theta)S(\rho)^{-1} \sim (a_2 \circ \exp(\pi H_{\sqrt{\hbar}}))(\theta, D_\theta).$$

Then Theorem 3 follows from an *inverse Egorov theorem*.

Finally we remark that this calculus can also be used to show the propagation properties of the *scattering wave front set* of Melrose, but we omit the detail here.

## References

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